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Università degli Studi di Padova

Small bodies

Geological and geo-structural Mapping

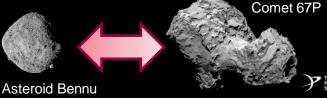
Matteo Massironi



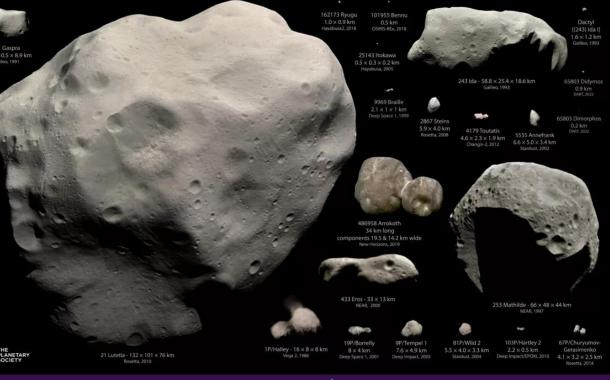
Small bodies of the Solar System

Comets are small bodies with highly elliptical orbits consisting largely of ices but also containing organics, dust and rocky material.

Asteroids largely consist of rocky material but also contain varying quantities of organics and volatiles.

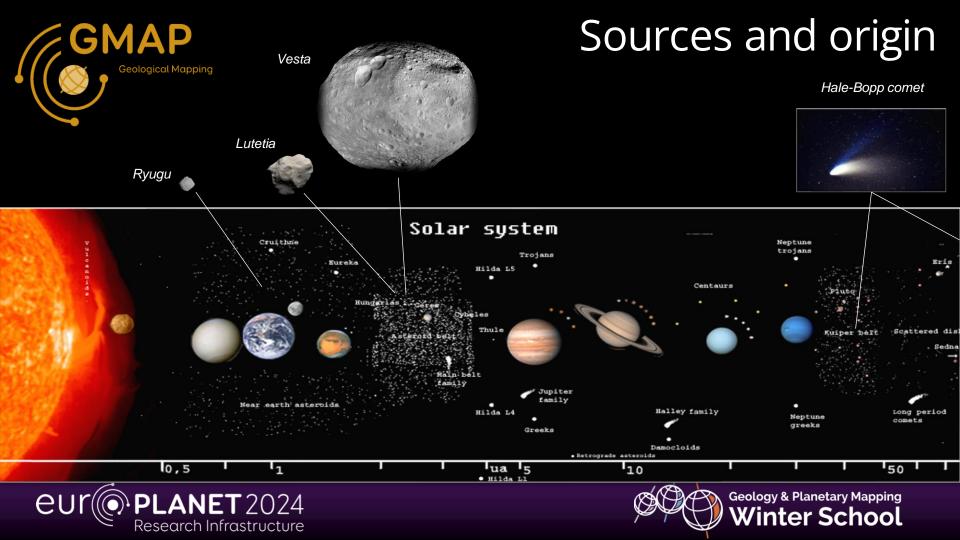








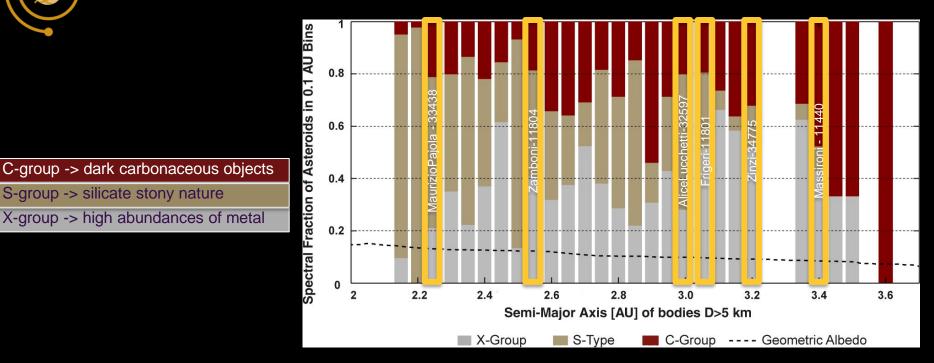
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Asteroid Classification



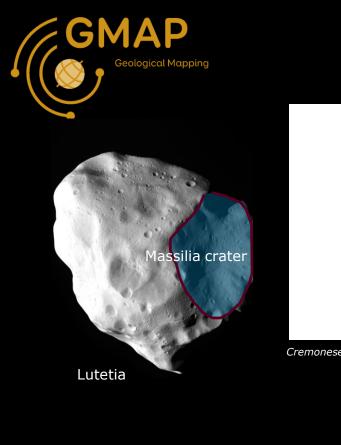
S-group -> silicate stony nature



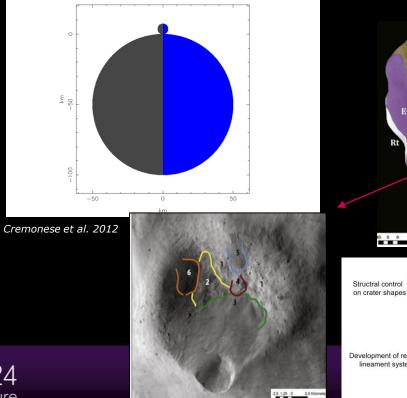
Schmedemann et al. in Rossi and Von Gasselt 2018





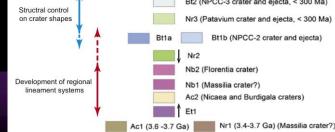


Main landforming processes on asteroids



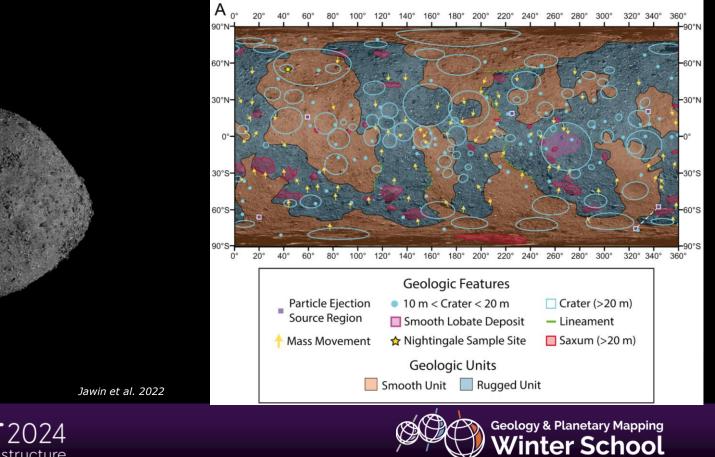
Ac2 Nb2 Btb Btb Btb Btb Nb1 Nb1 Nc1 Nc2 Massironi et al. 2012

Ac1







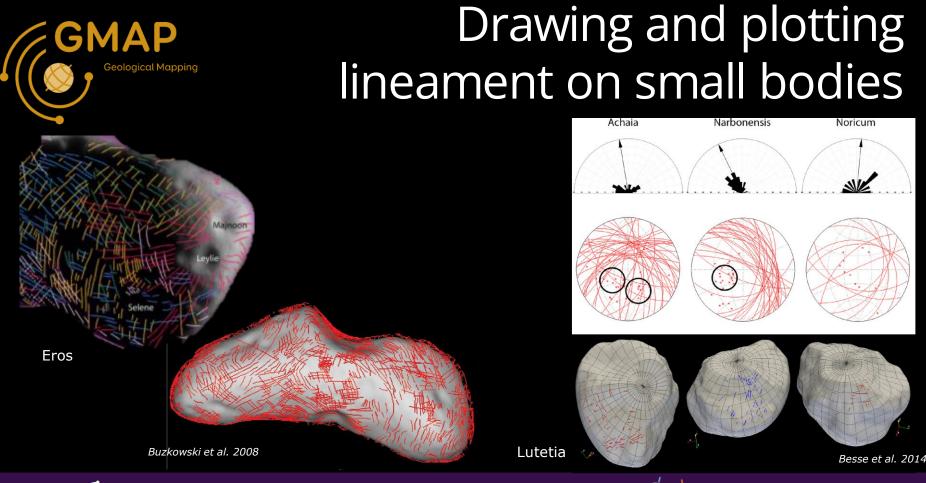




Bennu

GMAP

Geological Mapping



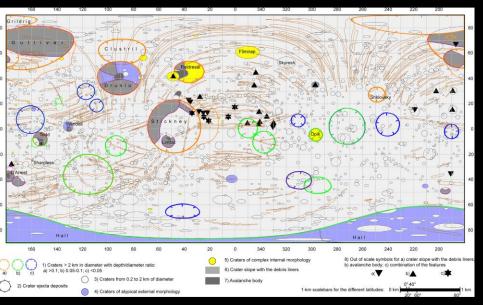




The Phobos case

Although a Martian moon it has the same geological landforms and

related mapping issues of any asteroid (small body)



Basilevsky et al. 2014



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a.3

b.3

Simioni et al. 2015

а

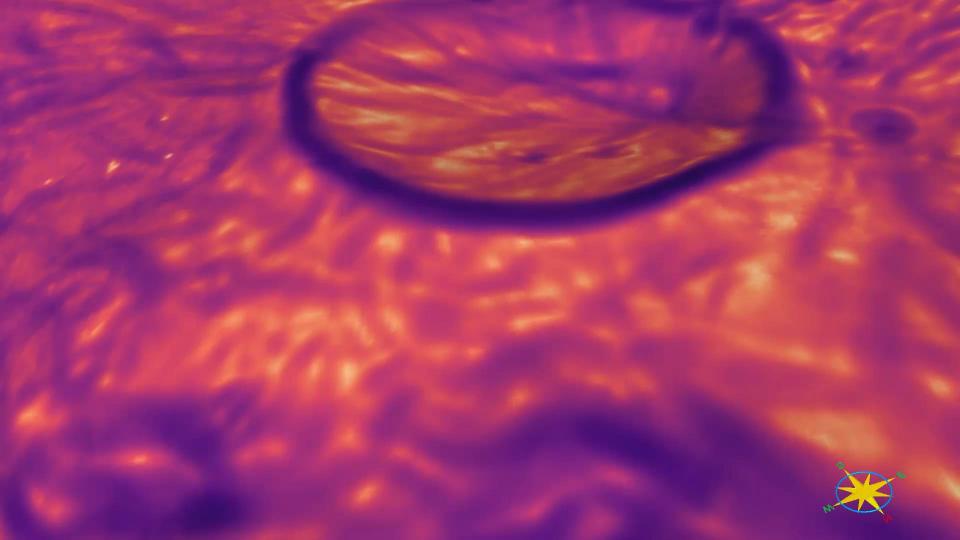
eu

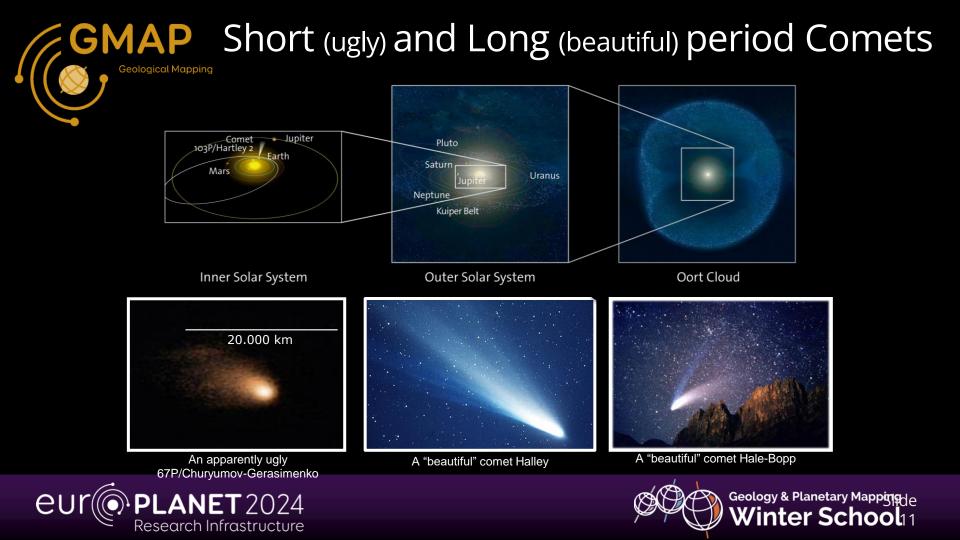
GMAP

Geological Mapping

2 km

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GMAP The Realms of Sublimation **Geological Mapping COMETARY NUCLEI** P/2005 JQ5 (Catalina) 45P/Honda-Mrkos-Paidušáková 1.4 km ~1.3 km Arecibo (2005) Arecibo (2017) 67P/Churyumov-Gerasimenko 81P/Wild 2 5 × 3 km 5.5 × 4.0 × 3.3 km Rosetta (2014) Stardust (2004) C/2013 A1 Siding Spring P/2016 BA14 Pan-STARRS ~0.4 km >1 km Mars Reconnaissance Orbiter (2014) Goldstone (2016) 73P/Schwassmann-Wachmann 3 Fragment C ~1 km Arecibo (2006) 73P/Schwassmann-Wachmann 3 9P/Tempel 1 Fragment B 7.5 × 5.0 km ~0.4 km Deep Impact (2005), Stardust-NExT (2011) Arecibo (2006) 1P/Halley 15.3 × 7.2 km Vega 1, Vega 2, Giotto (1986) 19P/Borrelly 8.0 × 3.2 km Deep Space 1 (2001) 8P/Tuttle 10 × 3.5 km 209P/LINEAR Arecibo (2008) 103P/Hartley 2 2.3 × 0.7 km 3 × 2.4 km Arecibe (2010), Deep Impact/EPOXI (2010) Arecibo (2014) 1 km 5 km

Image credits:

19//Eding: ESA / MPS (H. U. Kellen); 89//Tuttie: Arocibio Observatory / Mike Nalan / Daniel Macháče); 99/Tempof /: MASA/ JPL / UME / Daniel Macháče); 19//Edinzy/Morc-Genzimenko: ESA / Rosetta / NAVCAM; 759/Schwassmann-Wachmann & Arecibio Observatory / MSA / JPL / UME / Daniel Macháče); 19//Wild 2: NASA / JPL / Umbi / Macháče); 29//Edinzificoz: Arecibio Observatory / MSA / MSP / MSP / JPL / JPL







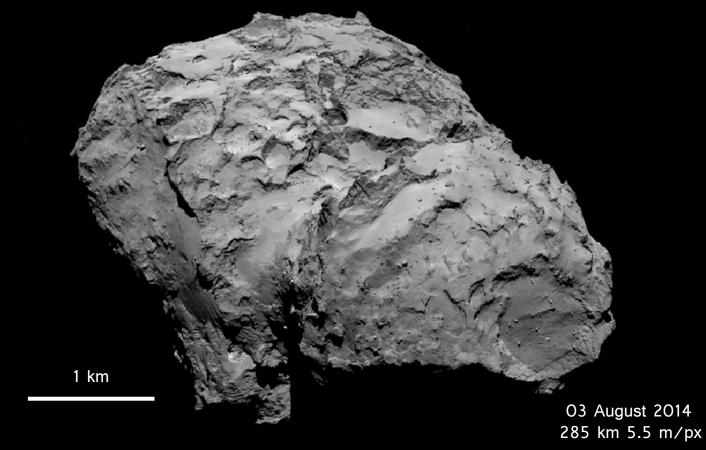
The Rosetta Journey to comet 67P (2 Mach 2004-6 August 2014)



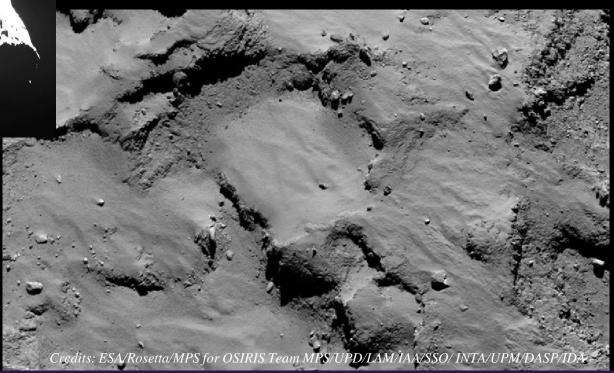
27/3, 3/4, 7/4, 14/4, 17/4, 20/4, 27/4, 30/4, 4/5 From 5 Million kms to 2 Million kms



03 August 2014 Rosetta meet 67P/CG



Cometary landforms: Outcrops, Airfalls and Taluses

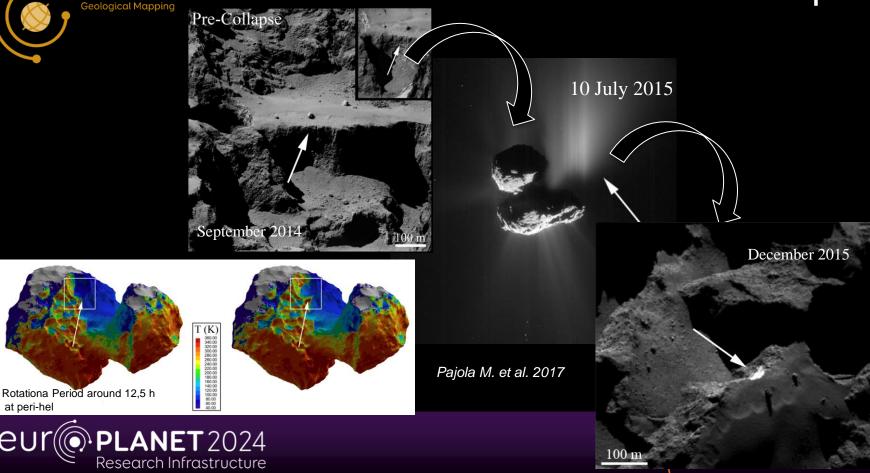








Outburst and Gravittional collapses

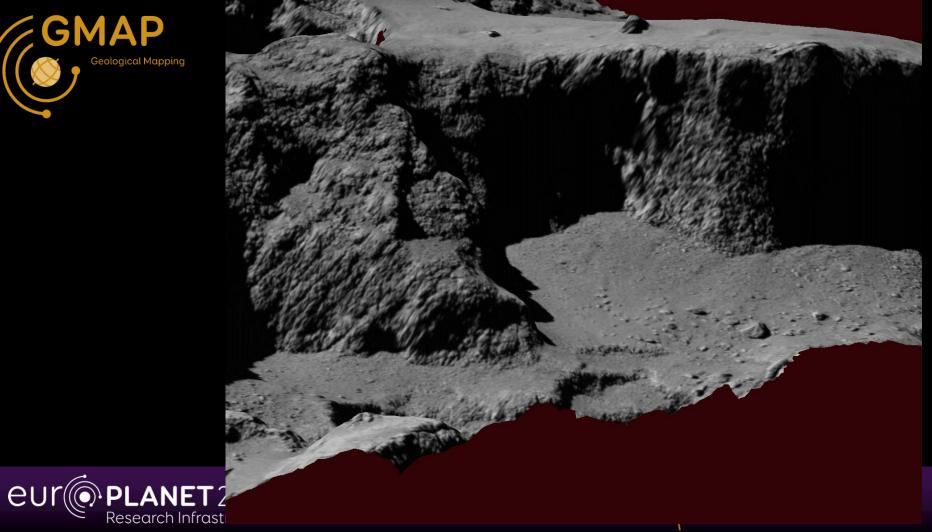


GMAP

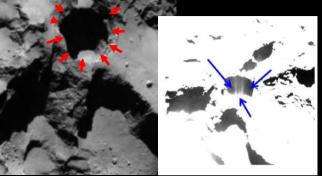
at peri-hel

eu



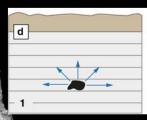


Active Pits



Vincent et al. 2015

2



500 m

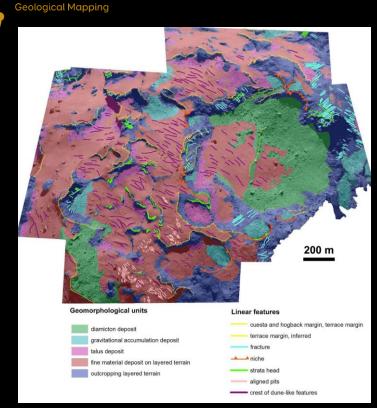


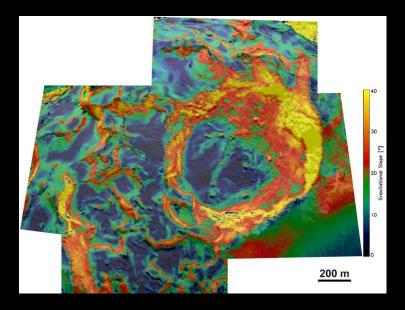




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Local geological maps





La Forgia et al. 2015

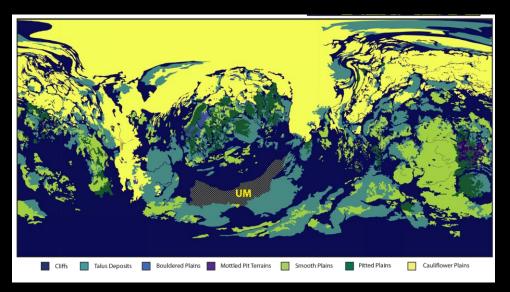


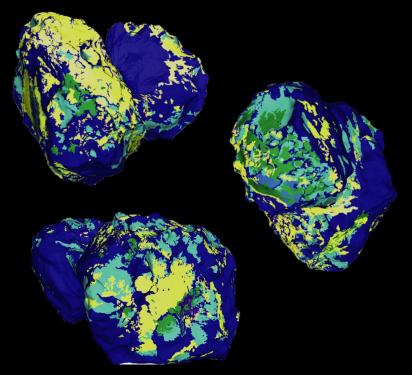
GMAP











Birch et al. 2017

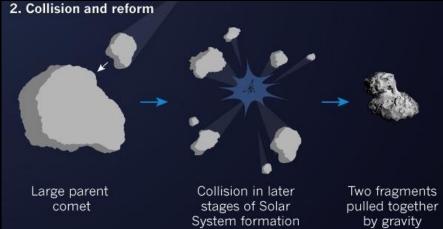






How the lobes formed?











Stair-case terraces







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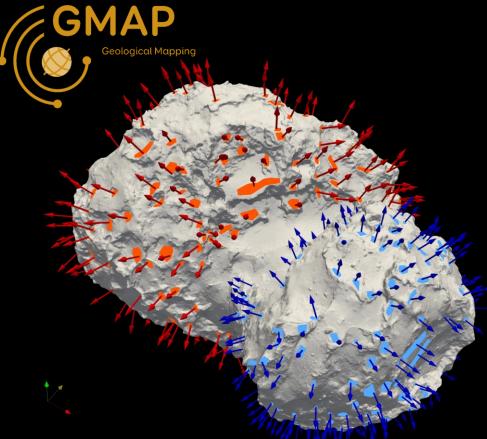
Regional layering



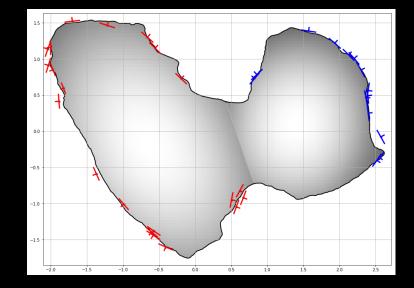


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500 m



Strata Measurements



Penasa et al. 2017





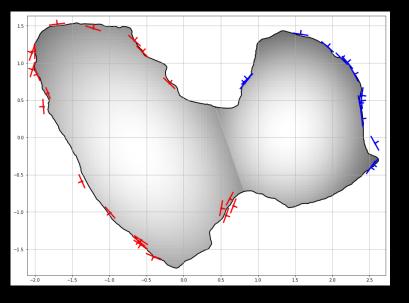


GMAP Geological Mapping Layer attitudes and gravity field Gravity



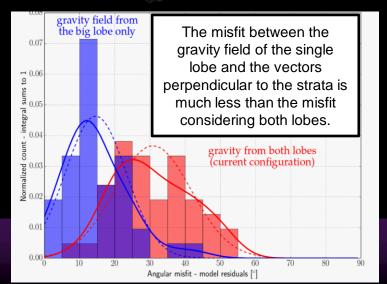
Gravity due to the BL

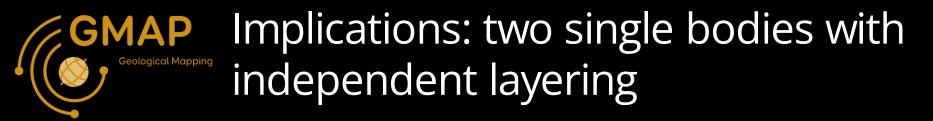
gravity

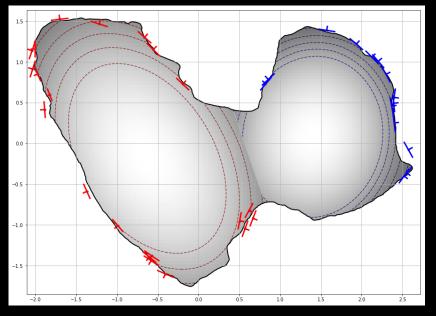


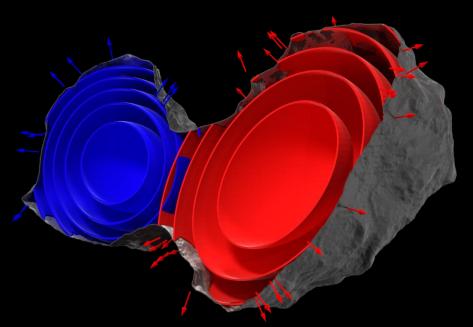
Penasa et al. 2017









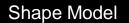


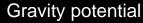
Penasa et al. 2017





Let's practice through VRGS









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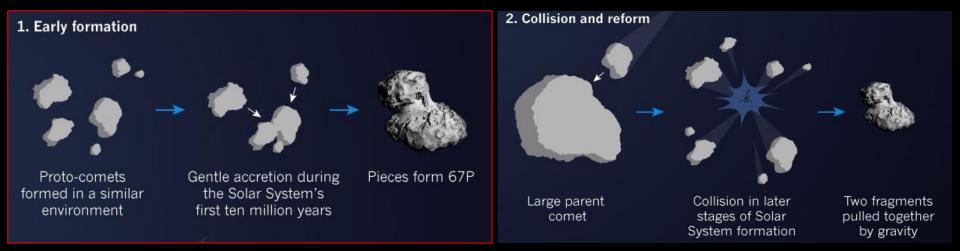


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Further Implications: an early aggregation

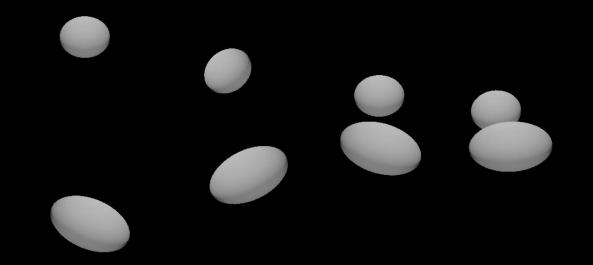








Further Implications: A gentle joining







Further implications: Avoiding catastrophic collisions

Nice Model Credit: Hal Levison, Southwest Research Institute, Boulder, Colorado

67P should have escaped catastrophic collisions during the Solar System evolution



GMAP

Geological Mapping



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